

The asymptotic structure of gravity at spatial infinity

Marc Henneaux

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and was identified first at null infinity by Bondi, Metzner and Sachs (“BMS group”)

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in the context of gravitational radiation

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The group of asymptotic symmetries of gravity in the asymptotically flat context

is infinite-dimensional

and was identified first at null infinity by Bondi, Metzner and Sachs (“BMS group”)

in the context of gravitational radiation

where null infinity is indeed a natural place to be.

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So, why should one study the asymptotic structure of gravity at spatial infinity in the asymptotically flat context?

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So, why should one study the asymptotic structure of gravity at spatial infinity in the asymptotically flat context?

Will give here three reasons.

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In order to better understand the role of the BMS group in the quantum theory, where physical states are usually defined on spacelike (Cauchy) hypersurfaces,

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First reason

In order to better understand the role of the BMS group in the quantum theory, where physical states are usually defined on spacelike (Cauchy) hypersurfaces,

it is important to unveil its action on spacelike (Cauchy) hypersurfaces, and thus, at spatial infinity.

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(First reason, continued)

This has been done recently in four spacetime dimensions, through a reconsideration of the boundary conditions at spatial infinity.

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This has been done recently in four spacetime dimensions, through a reconsideration of the boundary conditions at spatial infinity.

(Boundary conditions given originally at spatial infinity in the literature did not exhibit the BMS group.)

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(First reason, continued)

This has been done recently in four spacetime dimensions, through a reconsideration of the boundary conditions at spatial infinity.

(Boundary conditions given originally at spatial infinity in the literature did not exhibit the BMS group.)

New, consistent boundary conditions have been proposed, which are invariant under the full infinite-dimensional BMS group,

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(Boundary conditions given originally at spatial infinity in the literature did not exhibit the BMS group.)

New, consistent boundary conditions have been proposed, which are invariant under the full infinite-dimensional BMS group, providing a standard, non-trivial, canonical realization of the BMS symmetry.

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This has been done recently in four spacetime dimensions, through a reconsideration of the boundary conditions at spatial infinity.

(Boundary conditions given originally at spatial infinity in the literature did not exhibit the BMS group.)

New, consistent boundary conditions have been proposed, which are invariant under the full infinite-dimensional BMS group, providing a standard, non-trivial, canonical realization of the BMS symmetry.

This establishes also the important fact that the BMS symmetry is a symmetry of the theory and not just a symmetry at null infinity.

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Second reason

Another reason for investigating the asymptotic structure at spatial infinity is the need to understand the “matching” conditions of the fields and charges between \mathcal{I}_-^+ and \mathcal{I}_+^-

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Another reason for investigating the asymptotic structure at spatial infinity is the need to understand the “matching” conditions of the fields and charges between \mathcal{I}_-^+ and \mathcal{I}_+^- which clearly involves “going through” spatial infinity i^0 .

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Second reason

Another reason for investigating the asymptotic structure at spatial infinity is the need to understand the “matching” conditions of the fields and charges between \mathcal{I}_-^+ and \mathcal{I}_+^- which clearly involves “going through” spatial infinity i^0 .

This requires understanding the action of all the BMS supertranslations at spatial infinity.

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Second reason

Another reason for investigating the asymptotic structure at spatial infinity is the need to understand the “matching” conditions of the fields and charges between \mathcal{I}_-^+ and \mathcal{I}_+^- which clearly involves “going through” spatial infinity i^0 .

This requires understanding the action of all the BMS supertranslations at spatial infinity.

Connecting spatial infinity to the past of future null infinity (or the future of past null infinity) is actually a somewhat subtle question, because apparently “reasonable” Cauchy data might make null infinity not as smooth as one might have hoped, in the sense that the metric (and the Weyl tensor) develops logarithmic singularities.

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In five dimensions, the definition of null infinity is problematical (as it is in all odd spacetime dimensions).

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In five dimensions, the definition of null infinity is problematical (as it is in all odd spacetime dimensions).

But there exist soft theorems!

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Of which symmetries are these the Ward identities?

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Of which symmetries are these the Ward identities?

It turns out that while the answer to this question is not immediate at null infinity, the analysis at spatial infinity raises no conceptual problem

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and directly leads to the infinite dimensional symmetry “BMS₅ group”,

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Of which symmetries are these the Ward identities?

It turns out that while the answer to this question is not immediate at null infinity, the analysis at spatial infinity raises no conceptual problem

and directly leads to the infinite dimensional symmetry “BMS₅ group”,

the realization of which exhibits (somewhat unexpectedly) a very interesting nonlinear structure.

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The purpose of this talk is to provide the general ideas of the asymptotic analysis at spatial infinity in $D = 4$ and $D = 5$ spacetime dimensions.

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The purpose of this talk is to provide the general ideas of the asymptotic analysis at spatial infinity in $D = 4$ and $D = 5$ spacetime dimensions.

The study will be carried on spacelike hypersurfaces that are asymptotically flat hyperplanes, using Hamiltonian methods.

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(Work done in collaboration with Cédric Troessaert, Oscar Fuentealba, Sucheta Majumdar, Javier Matulich and Turmoli Neogi)

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(Work done in collaboration with Cédric Troessaert, Oscar Fuentealba, Sucheta Majumdar, Javier Matulich and Turmoli Neogi)

The analysis also provides an opportunity to develop general considerations on asymptotic symmetries.

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A central role in the analysis will be played by the gravitational action

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A central role in the analysis will be played by the gravitational action

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$$S[g_{ij}, \pi^{ij}, N, N^i] = \int dt \left\{ \int d^d x \left(\pi^{ij} \partial_t g_{ij} - N^i \mathcal{H}_i^{\text{grav}} - N \mathcal{H}^{\text{grav}} \right) - B_\infty \right\}$$

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where B_∞ is a boundary term at infinity and where

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where B_∞ is a boundary term at infinity and where

$$\mathcal{H}^{\text{grav}} = -\sqrt{g}R + \frac{1}{\sqrt{g}}(\pi^{ij}\pi_{ij} - \frac{1}{d-1}\pi^2) \approx 0, \quad \mathcal{H}_i^{\text{grav}} = -2\nabla_j \pi_i^j \approx 0.$$

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A central role in the analysis will be played by the gravitational action

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where B_∞ is a boundary term at infinity and where

$$\mathcal{H}^{\text{grav}} = -\sqrt{g}R + \frac{1}{\sqrt{g}}(\pi^{ij}\pi_{ij} - \frac{1}{d-1}\pi^2) \approx 0, \quad \mathcal{H}_i^{\text{grav}} = -2\nabla_j \pi_i^j \approx 0.$$

The definition of the theory is completed by providing boundary conditions on the dynamical variables (definition of phase space), which are assumed to make the action finite.

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A symmetry transformation is a phase space transformation (i.e., a transformation $(g_{ij}, \pi^{ij}) \rightarrow (g'_{ij}, \pi'^{ij})$ which preserves the boundary conditions)

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A symmetry transformation is a phase space transformation (i.e., a transformation $(g_{ij}, \pi^{ij}) \rightarrow (g'_{ij}, \pi'^{ij})$ which preserves the boundary conditions)

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A symmetry transformation preserves therefore in particular the symplectic form (“canonical transformation”)

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$$\Omega = \int d^d x d_V \pi^{ij} \wedge d_V g_{ij}$$

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exactly - and not up to surface terms.

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that leaves the action invariant up to surface integrals at the time boundaries.

A symmetry transformation preserves therefore in particular the symplectic form (“canonical transformation”)

$$\Omega = \int d^d x d_V \pi^{ij} \wedge d_V g_{ij}$$

exactly - and not up to surface terms.

Its canonical generator defines furthermore a constant of the motion.

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Hence it preserves not only the boundary conditions,

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Hence it preserves not only the boundary conditions, but it is also canonically generated.

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Hence it preserves not only the boundary conditions, but it is also canonically generated.

Its canonical generator is given by a bulk term proportional to the (first class) constraints plus a surface integral at infinity.

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When the constraints hold, it reduces to the surface term at infinity.

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Its canonical generator is given by a bulk term proportional to the (first class) constraints plus a surface integral at infinity.

When the constraints hold, it reduces to the surface term at infinity.

In gravity, all symmetries (with local action on the fields) are presumably asymptotic symmetries (no “rigid symmetry”, except perhaps non-local).

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Among the asymptotic symmetries, the “trivial” ones - also called “proper gauge transformations” - are the asymptotic symmetries purely generated by the constraints,

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Among the asymptotic symmetries, the “trivial” ones - also called “proper gauge transformations” - are the asymptotic symmetries purely generated by the constraints, **i.e., for which the surface terms vanish for all configurations obeying the boundary conditions.**

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i.e., for which the surface terms vanish for all configurations obeying the boundary conditions.

The asymptotic symmetries with non-vanishing generator are called “improper gauge symmetries”.

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The asymptotic symmetries with non-vanishing generator are called “improper gauge symmetries”.

They act non trivially on the physical states. They should not be factored out.

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The proper gauge symmetries form an ideal.

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The asymptotic symmetries with non-vanishing generator are called “improper gauge symmetries”.

They act non trivially on the physical states. They should not be factored out.

The proper gauge symmetries form an ideal.

The physical asymptotic symmetry algebra is the quotient of all the gauge transformations preserving the boundary conditions by the ideal of the proper gauge symmetries.

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Among the asymptotic symmetries, the “trivial” ones - also called “proper gauge transformations” - are the asymptotic symmetries purely generated by the constraints,

i.e., for which the surface terms vanish for all configurations obeying the boundary conditions.

The asymptotic symmetries with non-vanishing generator are called “improper gauge symmetries”.

They act non trivially on the physical states. They should not be factored out.

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(see R. Benguria, P. Cordero and C. Teitelboim, Nucl. Phys. B **122** (1977), 61-99)

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A prime example is given by the asymptotic charges themselves : when the asymptotic symmetry algebra is non-Abelian, they transform non-trivially.

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This is of course sensible on physical grounds.

E.g., a moving black hole is not physically the same as a black hole at rest (for an asymptotic inertial observer), even though the two situations differ by a diffeomorphism.

The moving black hole and the black hole at rest can be distinguished by the observable 4-momentum.

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They obviously depend on the boundary conditions.

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A striking example is higher spin gravity in three dimensions, described by a Chern-Simons theory.

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If the space is simply connected, there is no non-trivial observable invariant under all gauge transformations (the curvature vanishes), but non-trivial observables invariant only under proper gauge transformations exist.

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A striking example is higher spin gravity in three dimensions, described by a Chern-Simons theory.

If the space is simply connected, there is no non-trivial observable invariant under all gauge transformations (the curvature vanishes), but non-trivial observables invariant only under proper gauge transformations exist.

Depending on the boundary conditions, these form a W_3 or an inequivalent $W_3^{(2)}$ algebra.

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This is all right if the implicit gauge conditions freeze only proper gauge transformations.

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How do we know?

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Will discuss this question on the example of 4D gravity.

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$$h_{ij} \equiv g_{ij} - \delta_{ij} = \frac{\bar{h}_{ij}(\mathbf{n}^k)}{r} + O\left(\frac{1}{r^2}\right), \quad \bar{h}_{ij}(-\mathbf{n}^k) = \bar{h}_{ij}(\mathbf{n}^k)$$

and

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They involve strict parity conditions under the antipodal map $\mathbf{n}^k \rightarrow -\mathbf{n}^k$, where \mathbf{n}^k is the unit normal to the sphere at infinity

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The diffeomorphisms that preserve these parity conditions are the homogeneous Lorentz transformations (boosts and spatial rotations), the translations $\xi^\perp = a^0$, $\xi^i = a^i$ ($a^\mu = \text{constants}$)

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The diffeomorphisms that preserve these parity conditions are the homogeneous Lorentz transformations (boosts and spatial rotations), the translations $\xi^\perp = a^0$, $\xi^i = a^i$ ($a^\mu = \text{constants}$) and the “supertranslations” $\xi^\perp(\mathbf{n}^k)$ and $\xi^j(\mathbf{n}^k)$ of definite odd parity.

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With the above boundary conditions, the asymptotic symmetry group is the Poincaré group, a result at variance with the findings at null infinity.

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But perhaps this is due to an improper gauge fixing, freezing the non-trivial BMS supertranslations?

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But perhaps this is due to an improper gauge fixing, freezing the non-trivial BMS supertranslations?

This is indeed the case, because the standard BMS supertranslations have actually the opposite parity and thus are not allowed with the strict parity boundary conditions.

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One thus imposes

$$h_{ij} \equiv g_{ij} - \delta_{ij} = h_{ij}^{RT} + U_{ij}, \quad \pi^{ij} = \pi_{RT}^{ij} + V^{ij}$$

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$$h_{ij} \equiv g_{ij} - \delta_{ij} = h_{ij}^{RT} + U_{ij}, \quad \pi^{ij} = \pi_{RT}^{ij} + V^{ij}$$

U_{ij} and V^{ij} are the parity-twisted contributions that take the form of a gauge transformation (rewritten in Hamiltonian form).

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They are of the same order as the leading terms in h_{ij} and π^{ij} ($O(1/r)$ and $O(1/r^2)$ respectively) but have the opposite parity.

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They are of the same order as the leading terms in h_{ij} and π^{ij} ($O(1/r)$ and $O(1/r^2)$) respectively but have the opposite parity.

To preserve the boundary conditions, and for the Lorentz boosts to act canonically, one finds that U_{ij} is parametrized by an $O(1)$ odd function of the angles $\bar{U}(\mathbf{n}^k) = O(1) = -U(-\mathbf{n}^k)$ while V^{ij} is parametrized by an $O(1)$ even function of the angles $V(\mathbf{n}^k) = V(-\mathbf{n}^k)$.

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Do these relaxed parity conditions involving a twist lead to a consistent description

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The answer is affirmative and requires some work (even though the idea is elementary).

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One finds furthermore that the asymptotic symmetries are given by hypersurface deformations that behave asymptotically as

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One finds furthermore that the asymptotic symmetries are given by hypersurface deformations that behave asymptotically as

$$\xi = b_i x^i + T(\mathbf{n}) + O(r^{-1})$$

$$\xi^i = b^i_j x^j + W_i(\mathbf{n}) + O(r^{-1}), \quad b_{ij} = -b_{ji}, \quad W_i(\mathbf{n}) = \partial_i(rW(\mathbf{n})).$$

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where T is even and W is odd.

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The answer is affirmative and requires some work (even though the idea is elementary).

One finds furthermore that the asymptotic symmetries are given by hypersurface deformations that behave asymptotically as

$$\xi = b_i x^i + T(\mathbf{n}) + O(r^{-1})$$

$$\xi^i = b^i_j x^j + W_i(\mathbf{n}) + O(r^{-1}), \quad b_{ij} = -b_{ji}, \quad W_i(\mathbf{n}) = \partial_i(rW(\mathbf{n})).$$

where T is even and W is odd.

The terms $b_i x^i$ and $b^i_j x^j$ describe respectively boosts and spatial rotations.

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The zero mode of T and the first spherical harmonic component of W describe translations.

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In fact, the even function T and the odd function W combine to form a single arbitrary function of the angles, as in the null infinity description of the supertranslations.

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The symmetries are canonical transformations with generators

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$$P_\xi[g_{ij}, \pi^{ij}] = \int d^3x (\xi \mathcal{H} + \xi^i \mathcal{H}_i) + \mathcal{B}_\xi[g_{ij}, \pi^{ij}]$$

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The algebra of the generators can be easily verified to be the BMS algebra.

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From the BMS point of view, one can say that it is physically illegitimate to impose strict parity conditions as this requires improper gauge transformations,

which one cannot use in gauge fixing.

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define initial data that do not lead to log singularities as one goes to null infinity (at leading order).

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The opposite parity conditions would not fulfill this property and leading log singularities would be present at the critical surfaces corresponding to the past of future null infinity and the future of past null infinity.

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It is interesting to point out that the (twisted) parity conditions were introduced to make the Hamiltonian formalism satisfactory (finite action, canonical action of the Lorentz boosts).

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One gets a smoother null infinity as a (non-anticipated) bonus.

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But (???)

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The computations are cumbersome but the ideas are identical.

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As in 4 dimensions, one must include explicitly the improper gauge symmetries in the asymptotic form of the fields.

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A new feature is that improper gauge terms are not at the same order in $\frac{1}{r}$ as the Coulomb part,

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The computations are cumbersome but the ideas are identical.

As in 4 dimensions, one must include explicitly the improper gauge symmetries in the asymptotic form of the fields.

A new feature is that improper gauge terms are not at the same order in $\frac{1}{r}$ as the Coulomb part,

while in 4 dimensions, they are at the same order but distinguished by parity conditions.

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In 5 dimensions, the improper gauge terms are at order $\frac{1}{r}$ (for the metric), corresponding to a diffeomorphism parameter of order $\mathcal{O}(1)$, whereas “the rest” is at order r^{-2} (cf Schwarzschild in 5D)

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One has schematically :

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$$g_{ij} = \delta_{ij} + \mathcal{G}_{ij} + h_{ij}^{core}$$

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where $\mathcal{G}_{ij} = \mathcal{O}(\frac{1}{r})$ has the form of a diffeomorphism with vector field of order $\mathcal{O}(1)$ and where $h_{ij}^{core} = \mathcal{O}(r^{-2})$.

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Even though the ideas are identical, there are striking new features in 5 dimensions :

- The size of BMS_5 is bigger than expected. More precisely, the supertranslations depend on four functions of the angles.

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Even though the ideas are identical, there are striking new features in 5 dimensions :

- The size of BMS_5 is bigger than expected. More precisely, the supertranslations depend on four functions of the angles.
- Supertranslation charges (including the energy) acquire non-linear contributions.

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Even though the ideas are identical, there are striking new features in 5 dimensions :

- **The size of BMS_5 is bigger than expected. More precisely, the supertranslations depend on four functions of the angles.**
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- There are central charges among the different types of supertranslation generators.

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See O. Fuentealba, M. Henneaux, J. Matulich and C. Troessaert, e-Prints : 2111.09664 [hep-th] and 2206.04972 [hep-th]

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In terms of the canonical transformations generated by the charges, non-linearities mean that the commutator of the corresponding transformations is a transformation that takes the same form, but with coefficients that are functions of the fields through the charges.

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There is a great flexibility in the presentation of non-linear algebras since one can make non-linear redefinitions of the charges.

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The space generated by the symmetry charges is a Poisson manifold, about which much is known in the finite-dimensional.

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The space generated by the symmetry charges is a Poisson manifold, about which much is known in the finite-dimensional.

In particular, one can bring the Poisson bracket to a canonical form that generalizes the Darboux canonical form (“Weinstein splitting theorem” or “Darboux-Weinstein theorem”).

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In particular, one can bring the Poisson bracket to a canonical form that generalizes the Darboux canonical form (“Weinstein splitting theorem” or “Darboux-Weinstein theorem”).

The situation is much less understood in the infinite-dimensional case.

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One can bring the BMS_5 algebra to a form

where all the non-linearities in the Poincaré subalgebra have been absorbed through redefinitions, so that the Poisson brackets of the generators of boosts, spatial rotations and ordinary spacetime translations are the standard ones.

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where all the non-linearities in the Poincaré subalgebra have been absorbed through redefinitions, so that the Poisson brackets of the generators of boosts, spatial rotations and ordinary spacetime translations are the standard ones.

The non-linearities occur only in the brackets of the boosts with some of the proper supertranslations.

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One can bring the BMS_5 algebra to a form

where all the non-linearities in the Poincaré subalgebra have been absorbed through redefinitions, so that the Poisson brackets of the generators of boosts, spatial rotations and ordinary spacetime translations are the standard ones.

The non-linearities occur only in the brackets of the boosts with some of the proper supertranslations.

The supertranslations form an abelian subalgebra, with a central extension (that vanishes for the spacetime translations).

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Contrary to the familiar ADM expression for the energy in 4D, the energy in 5D (with our relaxed boundary conditions) acquires nonlinear contributions.

Explicitly,

$$E = 2 \oint_{S_\infty^3} d^3x \sqrt{\bar{\gamma}} \left[(1/2) \bar{h}_A^A + D_A \bar{\lambda}^A + 3\bar{\lambda} - (1/8) \theta_{AB} \theta^{AB} + K_A K^A \right]$$

where

$$g_{rr} = 1 + \frac{2\bar{\lambda}}{r^2} + \frac{h_{rr}^{(2)}}{r^3} + \mathcal{O}(r^{-4}), \quad (5.1)$$

$$g_{rA} = \frac{\bar{\lambda}_A}{r} + \frac{h_{rA}^{(2)}}{r^2} + \mathcal{O}(r^{-3}), \quad (5.2)$$

$$g_{AB} = r^2 \bar{g}_{AB} + r \theta_{AB} + \bar{h}_{AB} + \frac{h_{AB}^{(2)}}{r} + \mathcal{O}(r^{-2}). \quad (5.3)$$

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as it follows by applying standard canonical methods. (K_A comes from the momenta and non-linear redefinitions but plays no role below.)

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For instance if we perform the coordinate transformation $r = \rho + c$ on the flat metric $dr^2 + r^2 d\Omega^2$

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(first term : ADM contribution ; second term : nonlinear contribution)

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As one increases the dimension, the analysis becomes more and more technically intricate

because the gap between the pure (improper) diffeomorphism piece in the expansion of the fields (generated by $O(1)$ vector fields) and the “Coulomb” piece widens by one power of $1/r$ as one increases the dimension by one.

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Non linearities (of increasing order) then proliferate.

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Non linearities (of increasing order) then proliferate.

(Note that these non-linearities are not seen in current null infinity treatments which linearize the theory at infinity)

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Non linearities (of increasing order) then proliferate.

(Note that these non-linearities are not seen in current null infinity treatments which linearize the theory at infinity)

Preliminary analysis indicates that the size of the BMS group does not increase because the new terms in the diffeomorphism generators define proper gauge transformations.

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Such a term cannot be set to zero,

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One can then construct a completely consistent canonical formulation of the BMS symmetry,

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One can then construct a completely consistent canonical formulation of the BMS symmetry,

not only in four spacetime dimensions where the description is in complete agreement with the null infinity results (providing furthermore new light on the matching conditions),

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One can then construct a completely consistent canonical formulation of the BMS symmetry,

not only in four spacetime dimensions where the description is in complete agreement with the null infinity results (providing furthermore new light on the matching conditions),

but also in five dimensions where there is currently no null infinity description.

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Similar non-linear asymptotic algebras occur in 3D and in supergravity.

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Deriving the implications of the BMS_5 symmetry requires a better understanding of the physically relevant representations of the BMS symmetry (in non-separable Hilbert spaces?)

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Deriving the implications of the BMS_5 symmetry requires a better understanding of the physically relevant representations of the BMS symmetry (in non-separable Hilbert spaces?)

and the implications of the non-linearities in the algebra.

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THANK YOU!